

# From Strings to $AdS_4$ -Black holes

by Marco Rabbiosi

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# Why quantum theory of Gravity?

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- Electromagnetic, weak and strong  $\rightarrow$  **QFT** (Standard Model)
- Gravity  $\rightarrow$  **Differential Geometry** (General Relativity)

- These are two very *different mathematical frameworks*...  
..but a common belief is the existence of a *fundamental microscopic theory* able to describe all together.
- *Canonical quantization of GR fails* because of problems with unitarity and renormalizability...  
...however for a complete understanding of *Black holes* and *Big Bang* we need Quantum Gravity: they exhibit a *singular behaviour in GR*, curvature scalars diverge.
- Characteristic length:  $l_{ph} = \sqrt{\frac{\hbar G}{c^3}} \sim 10^{-35} m \rightarrow$  direct experimental data inaccessible...but on the theoretical side...

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- Enlargement of  $ISO(3,1) \rightarrow$  **Coleman-Mandula No-go theorem.**

4d SuperPoincaré algebra  $\mathfrak{s}_{\mathcal{N}} = \{P_\mu, J_{\mu\nu}, Q_\alpha^i = (Q_A^i, \bar{Q}_{Ai})^t, T_r\}$

$$\{Q_A^i, \bar{Q}_{Bj}\} = -2i(\sigma^\mu)_{AB} \delta_j^i P_\mu, \quad \{Q_A^i, Q_B^j\} = \varepsilon_{AB} (U^{ij} + iV^{ij}),$$

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plus complex conjugate relations and the usual Poincaré algebra.

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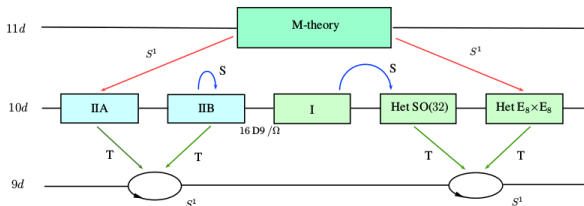
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- **Gauging supersymmetry** requires diffeomorphism invariance, ...the symmetry of GR!...we obtain **Supergravity!**
- Moreover, the **quantum spectrum of a string** always has a massless spin 2 particle...**again Gravity!**

## These ingredients lead to *M-theory*

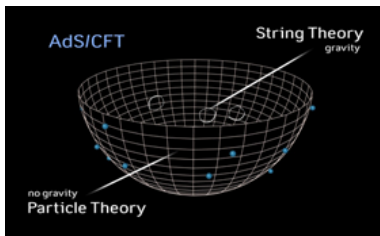
- $d = 11$  supersymmetric theory of **strings and branes**.
- Beyond an ordinary QFT: a **net of duality between superstring theories** that describe it in different limits.



- Interesting to study the *low energy limit* ( $\alpha' \rightarrow 0$ ) of these theories, the *massless part of the spectrum*.
- The *effective field theory* results to be a *Supergravity* theory in  $d = 10, 11$ !
- *Compactifications* means they admit solutions like  $\mathcal{C}_s \times \mathcal{M}_{d-s}$ .
- They *link the different Supergravities* and their deformations *in all the dimensions*!
- In particular  $\mathcal{N} = 2$   $d = 4$  *gauged supergravity* comes from compactifications on a flux background.
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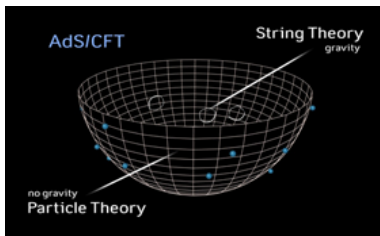
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$$AdS_4 \times S^7 / \mathbb{Z}_k \iff d = 3 \mathcal{N} = 6 \text{ ABJM } U(N)_k \times U(N)_{-k}$$

- *Asintotically*  $AdS_4$  *black hole* solutions could be *dual to some deformations of ABJM*.
- Lifting to *vacuas of M-theory*...  $M2$ -branes wrapping a Riemann surface.
- Study of *microstates of black holes*.
- *Integrability properties* of classical GR in  $d = 4$ .



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# $\mathcal{N} = 2$ $d = 4$ abelian gauged supergravity

- The **bosonic part** of the action (fermionic configuration to zero is a consistent truncation) **reads**

$$\begin{aligned} \mathcal{L} = & \sqrt{-g} \left( \frac{R}{2} - h_{uv} \nabla_\mu q^u \nabla^\mu q^v - g_{ij} \partial_\mu z^i \partial^\mu \bar{z}^{\bar{j}} + \frac{1}{4} l_{\Lambda\Sigma} H^{\Lambda\mu\nu} H_{\mu\nu}^\Sigma \right. \\ & + \frac{R_{\Lambda\Sigma} H_{\mu\nu}^\Lambda \varepsilon^{\mu\nu\rho\sigma} H_{\rho\sigma}^\Sigma}{8\sqrt{-g}} - (4h_{uv} k_\Lambda^u k_\Sigma^v X^\Lambda \bar{X}^\Sigma + (f_i^\Lambda g^{\bar{i}\bar{j}} \bar{f}_{\bar{j}}^\Sigma - 3X^\Lambda \bar{X}^\Sigma) P_\Lambda^x P_\Sigma^x) \\ & \left. - \frac{1}{4} \frac{\varepsilon^{\mu\nu\rho\sigma}}{\sqrt{-g}} \Theta^{\Lambda a} B_{\mu\nu a} \partial_\rho A_{\sigma\Lambda} + \frac{1}{32\sqrt{-g}} \Theta^{\Lambda a} \Theta_\Lambda^b \varepsilon^{\mu\nu\rho\sigma} B_{\mu\nu a} B_{\rho\sigma b} \right), \end{aligned}$$

- Where the **covariant derivative** of the hyperscalars is

$$\nabla_\mu q^u = \partial_\mu q^u + A_\mu^\Lambda k_\Lambda^u - A_{\Lambda\mu} k^{\Lambda u},$$

and  $H_{\mu\nu}^\Lambda = F_{\mu\nu}^\Lambda + \frac{1}{2} \Theta^{\Lambda a} B_{\mu\nu a}$  are **modified field strengths** (necessary for having also magnetically charged hyperscalars).

- We choose the **more simple possible ansatz** for the **metric**

$$ds^2 = -e^{2U(r)} dt^2 + e^{-2U(r)} dr^2 + e^{2\psi(r)-2U(r)} d\Omega_l^2,$$

where

$$f_l(\theta) = \frac{1}{\sqrt{l}} \sin(\sqrt{l}\theta) = \begin{cases} \sin \theta, & l = 1, \\ \sinh \theta, & l = -1, \end{cases}$$

and for the **gauge fields**

$$A^\Lambda = A_t^\Lambda dt - lp^\Lambda f_l'(\theta) d\phi, \quad A_\Lambda = A_{\Lambda t} dt - lq_\Lambda f_l'(\theta) d\phi,$$

$$B^\Lambda = 2lp'^\Lambda f_l'(\theta) dr \wedge d\phi, \quad B_\Lambda = -2lq'_\Lambda f_l'(\theta) dr \wedge d\phi$$

$$H_{tr}^\Lambda = e^{2U-2\psi} l^\Lambda \Sigma (R_{\Lambda\Gamma} p^\Gamma - q_\Sigma), \quad H_{\theta\phi}^\Lambda = p^\Lambda f_l(\theta).$$

- Now **all the fields** defining this ansatz **depend only on the radial coordinate  $r$ !**



- The substitution of this ansatz in the e.o.m. shows the possibility to derive these equations from a finite dimensional ***dynamical effective system***

$$S_{\text{eff}} = \int dr [e^{2\psi} (U'^2 - \psi'^2 + h_{uv} q'^u q'^v + g_{i\bar{j}} z'^i \bar{z}'^{\bar{j}} + \frac{1}{4} e^{4U-4\psi} \mathcal{Q}' \mathcal{H}^{-1} \mathcal{Q}') - V_{\text{eff}}],$$

with

$$V_{\text{eff}} = - \left( e^{2U-2\psi} V_{BH} + e^{2\psi-2U} V_g - l \right).$$

ensuring the constraints  $H_{\text{eff}} = 0$  and  $p^\Lambda k_\Lambda^u - q_\Lambda k^{u\Lambda} = 0$

- The e.o.m. now are switched *from PDE to ODE...*
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- For a *n-dimensional dynamical system* like

$$S = \int dr \left( \frac{1}{2} g_{ab} \Phi'^a \Phi'^b - V(\Phi^a) \right),$$

with the condition  $H = 0$ , Hamilton-Jacobi equation for the principal function  $W = W(\Phi^a; \alpha^1, \dots, \alpha^n)$  reads

$$\frac{1}{2} g^{ab} \partial_a W \partial_b W = -V(\Phi^a).$$

- In general, too difficult to find the complete integral...easier to have a particular solution!
- In this case we have not solved completely the system, however we can *rewrite the action as sum of square*

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Applying this technique to  $S_{\text{eff}}$  showed before, we find:

$$\begin{aligned}U' &= \varepsilon(e^{U-2\Psi} \text{Re} \tilde{\mathcal{L}} + i e^{-U} \text{Im} \tilde{\mathcal{L}}), \\ \Psi' &= 2 i \varepsilon e^{-U} \text{Im} \tilde{\mathcal{L}}, \\ z'^i &= \varepsilon e^{i\alpha} g^{i\bar{j}} \left( e^{U-2\Psi} \bar{D}_{\bar{j}} \tilde{\mathcal{L}} - i e^{-U} \bar{D}_{\bar{j}} \tilde{\mathcal{L}} \right), \\ q'^u &= -2 i \varepsilon h^{\mu\nu} e^{-U} \text{Im}(e^{-i\alpha} \partial_\nu \mathcal{L}), \\ \mathcal{Q}' &= \varepsilon 4 e^{2\Psi-3U} \mathcal{H} \Omega \text{Re} \tilde{\mathcal{V}},\end{aligned}$$

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Where the *principal Hamilton-Jacobi* function reads

$$W = e^U | \mathcal{L} + i e^{2\Psi-2U} \mathcal{Q}^x \mathcal{W}^x |,$$

- We have generalized the result of BPS analysis of Halmagyi, Petrini, Zaffaroni [arXiv : 1305.0730v3[hep-th]] and Dall'Agata, Gnechchi [arXiv : 1012.3756v1[hep-th]].

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# Summary and Outlook

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- The next step is to solve the equations in some *particular models and further extend the result* to a large class of black holes: non-extremal, rotating and NUT-charged...
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